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Accession number:20122015026298

Title:High-resolution absorption spectroscopy of the OH ²Π <inf>3/2</inf> ground state line

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Source title: Astronomy and Astrophysics

Abbreviated source title:Astron. Astrophys.

Volume:542

Issue date:2012

Publication year:2012

Article number:L7

Language:English

ISSN:00046361

E-ISSN:14320746

CODEN:AAEJAF

Document type: Journal article (JA)

Publisher: EDP Sciences, 17 Avenue du Hoggar - BP 112, Les Ulis Cedex A, F-91944, France

Abstract: The chemical composition of the interstellar medium is determined by gas phase chemistry, assisted by grain surface reactions, and by shock chemistry. The aim of this study is to measure the abundance of the hydroxyl radical (OH) in diffuse spiral arm clouds as a contribution to our understanding of the underlying network of chemical reactions. Owing to their high critical density, the ground states of light hydrides provide a tool to directly estimate column densities by means of absorption spectroscopy against bright background sources. We observed onboard the SOFIA observatory the $\langle \sup \rangle 2 \langle \sup \rangle \Π \langle \inf \rangle 3/2 \langle \inf \rangle$, J = 5/2 &rarr 3/2 2.5 THz line of ground-state OH in the diffuse clouds of the Carina-Sagittarius spiral arm. OH column densities in the spiral arm clouds along the sightlines to W49N, W51 and G34.26+0.15 were found to be of the order of 10 $\langle \sup \rangle 14 \langle \sup \rangle -2 \langle \sup \rangle -2 \langle \sup \rangle$, which is comparable to that of H $\langle \inf \rangle 2 \langle \inf \rangle O$. The absorption spectra of both species have similar velocity components, and the ratio of the derived H $\langle \inf \rangle 2 \langle \inf \rangle O$ to OH column densities ranges from 0.3 to 1.0. In W49N we also detected the corresponding line of $\langle \sup \rangle 18 \langle \sup \rangle OH$. © 2012 ESO.

Number of references:32

Main heading:Ground state

Controlled terms:Surface reactions

Uncontrolled terms:Chemical compositions - Column density - Critical density - Gas phase chemistry - High resolution - Hydroxyl radicals - Interstellar mediums - ISM : clouds - ISM: lines and bands - ISM: molecules - States of light - Underlying networks - Velocity components Classification code:802.2 Chemical Reactions - 932 High Energy Physics; Nuclear Physics;

Plasma Physics - 933 Solid State Physics DOI:10.1051/0004-6361/201218915 Database:Compendex Compilation and indexing terms, Copyright 2012 Elsevier Inc.